

*“Workshop on Atomic Effects in Nuclear Excitation and Decay”
June 15-19, 2009, ECT*, Trento, Italy*

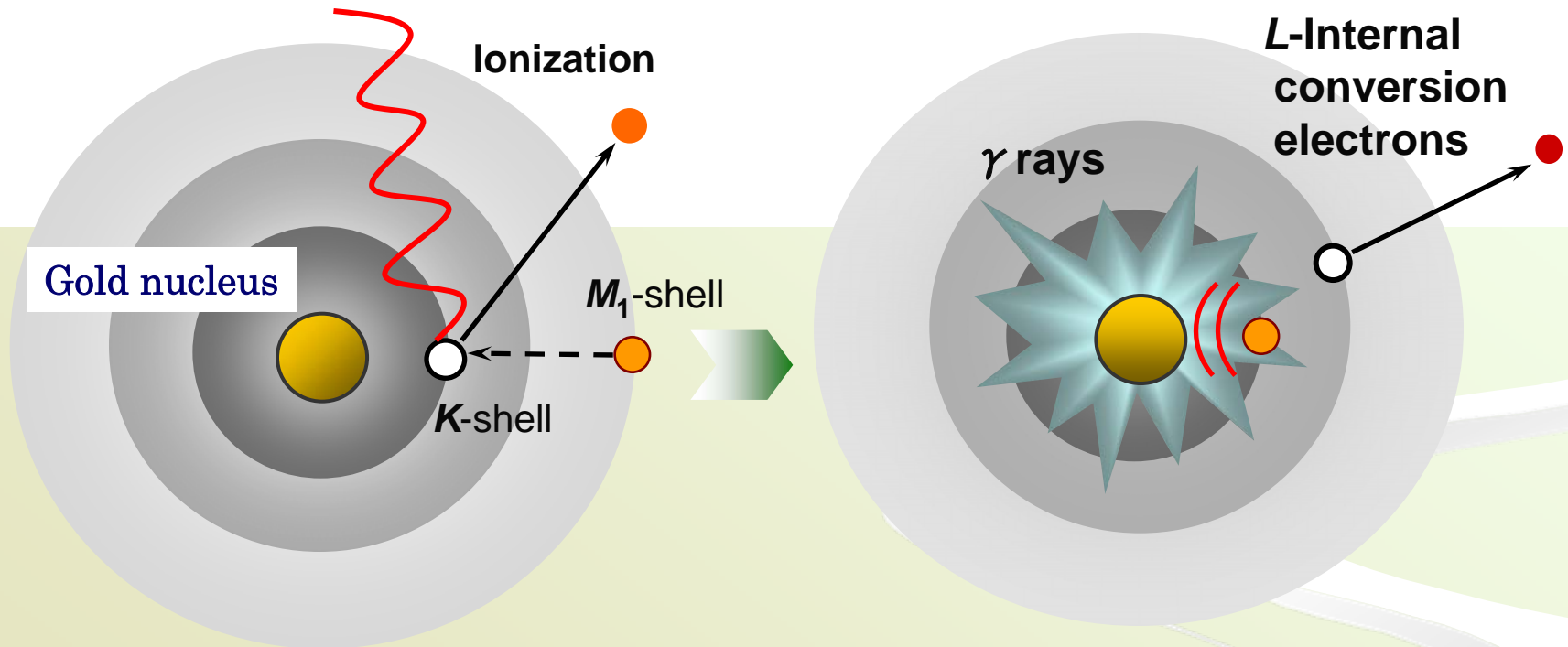
*Observation of
Nuclear Excitation by Electron Transition
using Synchrotron X-rays*

Shunji Kishimoto

*Photon Factory,
Institute of Materials Structure Science
High Energy Accelerator Research Organization (KEK),
Japan*

1. Nuclear Excitation by Electron Transition (NEET)

Ex. ^{197}Au , Phys. Rev. Lett. 85, 1831(2000)

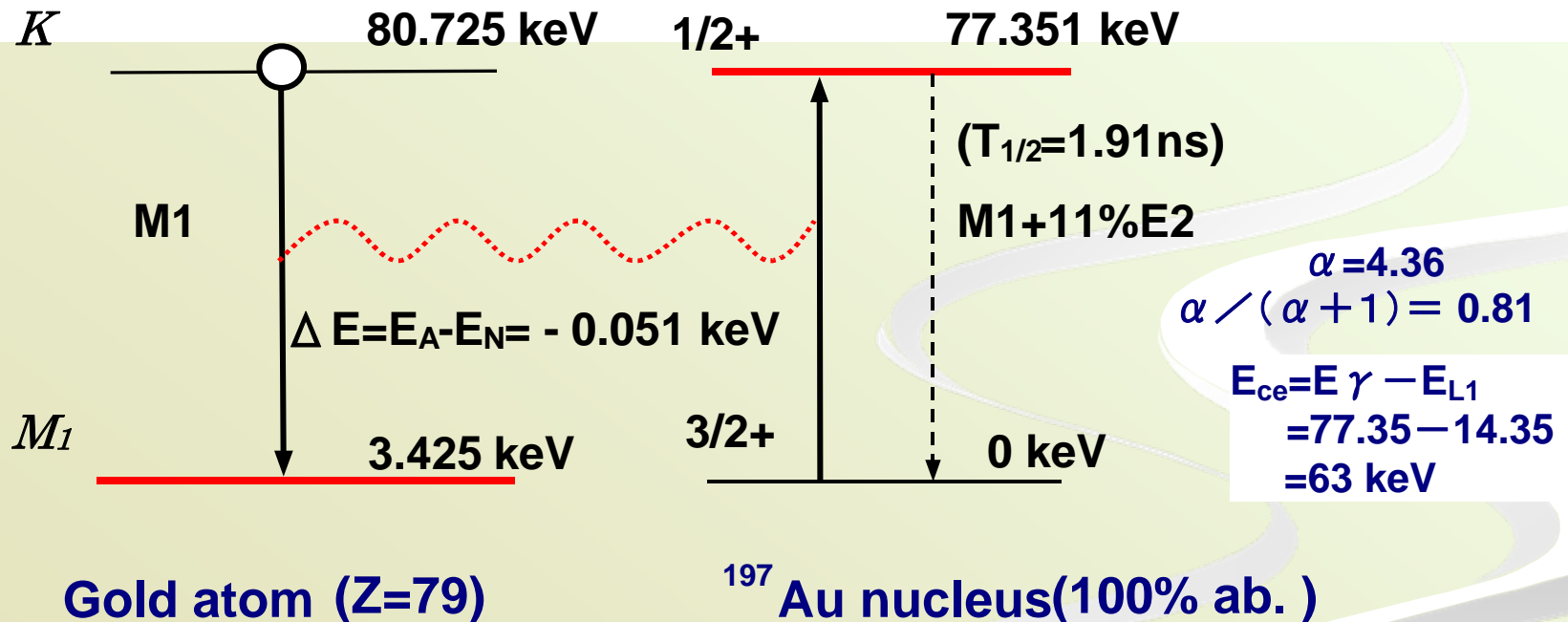


K-holes are produced by photoionization, and filled by an atomic transition from an outer orbit (Ex. M_1).

The nucleus is then excited and followed by emitting radiation with a lifetime of the excited level.

NEET is observable if the energies of **Atomic transition** and of **Nuclear transition** are close and the transition type is common to both.

Ex. Au-197



NEET Probability :

$$P_{\text{NEET}} = (\text{Nuclear Excitation}) / (\text{Core-electron hole Production})$$

M. Morita, Prog. Theor. Phys. 49, 1574 (1973) .

Nucleus Nat. ab.	E_{ion} (keV)	E_{M} (keV)	E_{N} (keV)	$T_{1/2}$ (ns)	ΔE (keV) $E_{\text{A}}-E_{\text{N}}$	P_{NEET} (cal.) (exp.)
^{197}Au (Z=79) 100%	80.725 (K)	3.425(M ₁)	77.351	1.91	-0.051	^A 3.8×10^{-8} ^B $(5.0 \pm 0.6) \times 10^{-8}$
^{193}Ir (Z=77) 62.7%	76.111 (K)	3.174(M ₁)	73.044	6.09	-0.107	^C 2.0×10^{-9} ^D $(2.8 \pm 0.4) \times 10^{-9}$

^A M.R.Harston, Nucl. Phys. A690(2001) 447.

^B S.Kishimoto et. al., Phys. Rev. Lett. 85(2000) 1831.
Phys.Rev.C74(2006),031301(R).

^C E.V.Tkalya, Phys. Rev. A75 (2005) 022509.

^D S.Kishimoto et. al., Nucl. Phys. A748(2005) 3.

2. How can we observe NEET ? - Synchrotron X-ray experiments

NEET: Rare events !

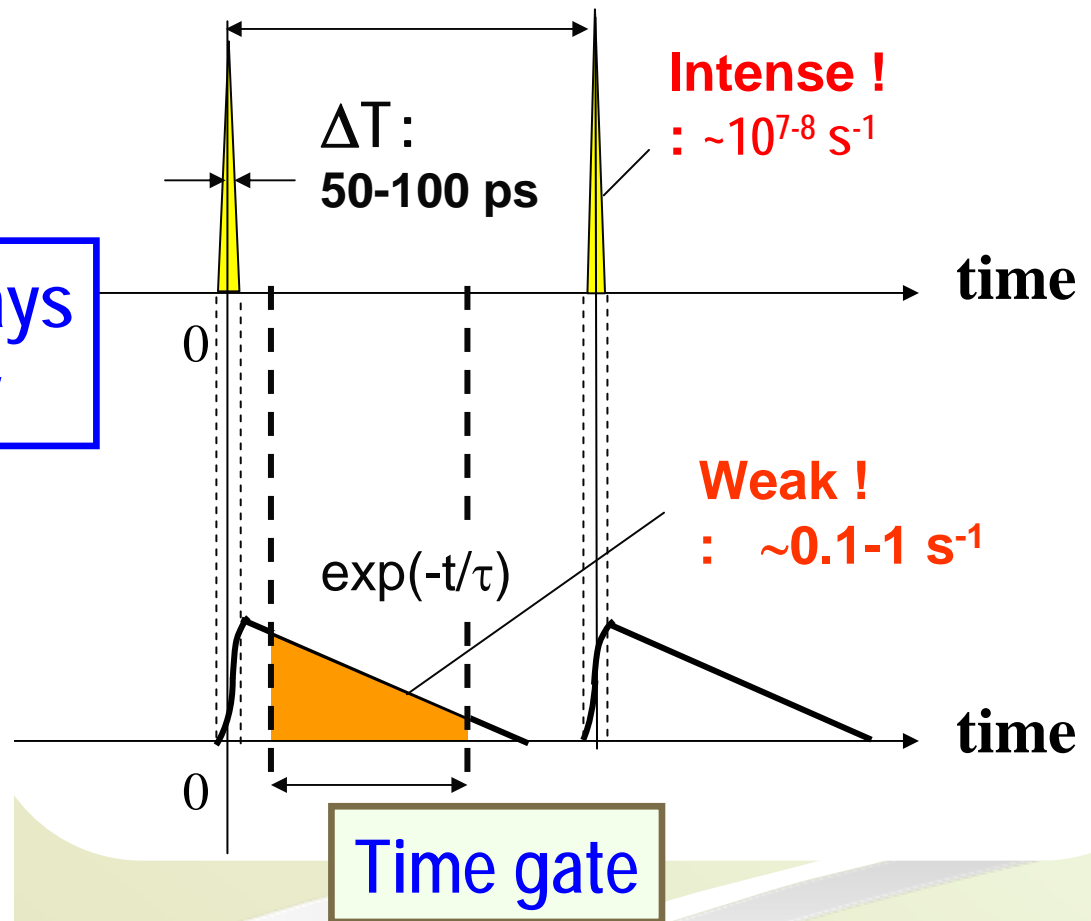
Prompt radiation by electronic scattering is more intense.

T: 10-1000 ns (a single or several-bunch mode)

Pulsed synchrotron X-rays
and Time Spectroscopy

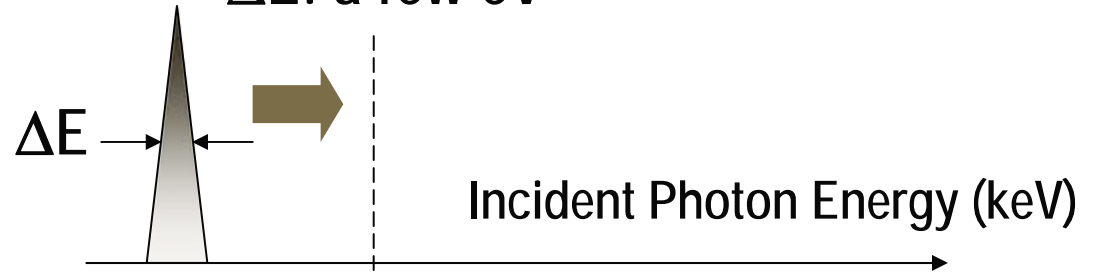


*Nuclear-decay Events :
can be separated
from the prompt
electronic scattering.*



NEET can be confirmed by choosing the incident photon energy where decaying radiation is detected.

Monochromatic X-ray beam
 ΔE : a few eV



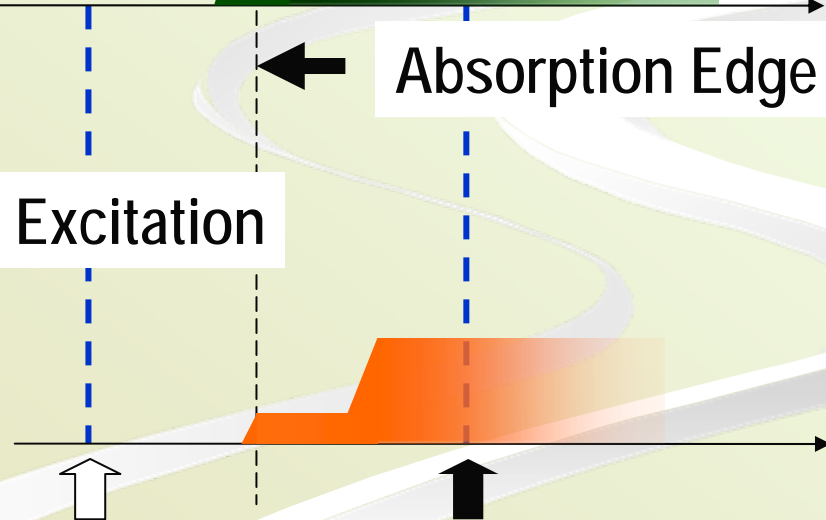
X-ray Absorption



Absorption Edge

Synchrotron X-ray is tunable in energy:

Nuclear Excitation



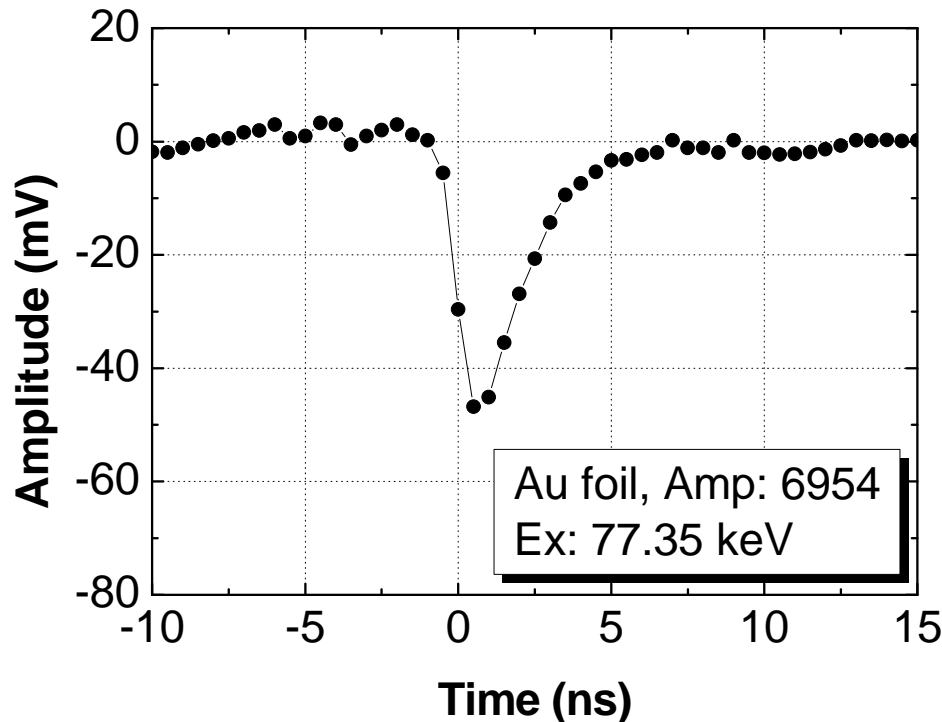
Detection of delayed electrons under the intense prompt radiation by Silicon Avalanche Diode

Si-AD : Hamamatsu S5344LC

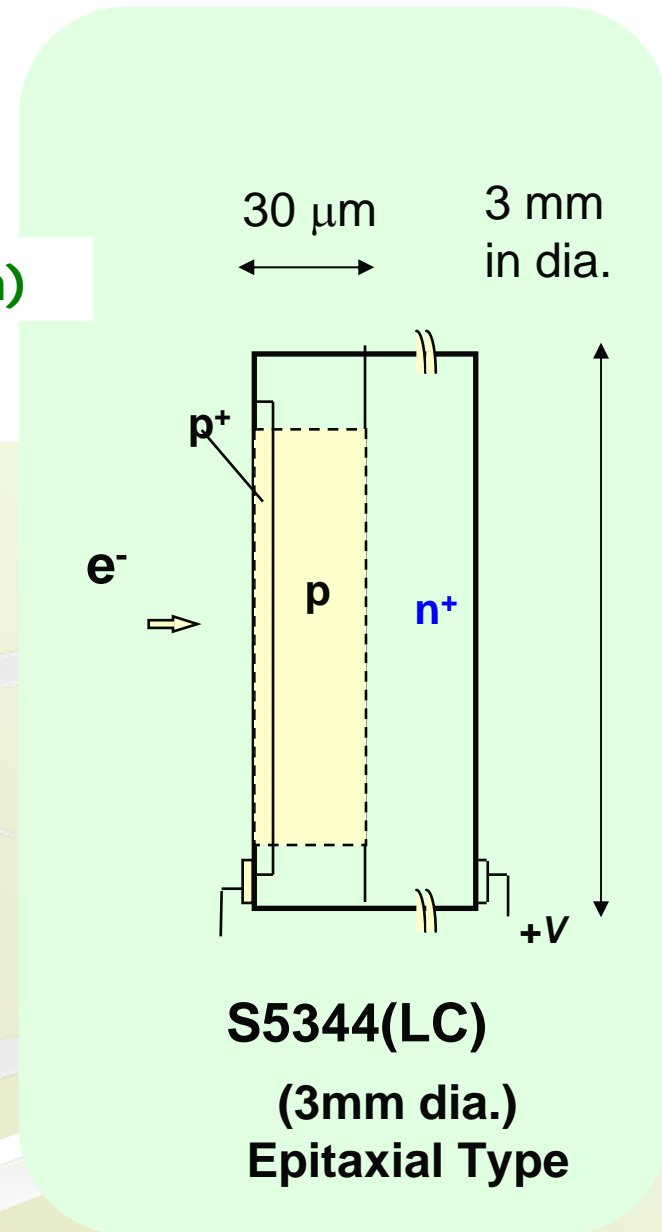
30 μm thick (L_1 conv. $E=63\text{keV}$, $\lambda_e \sim 24\mu\text{m}$)

➔ **a thin SiO_2 layer : 15-20nm (cf. std. 130nm)**

Low efficiency for high energy X rays ($< 0.1\%$)

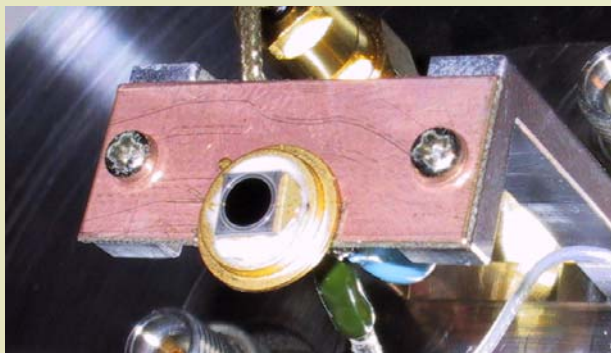
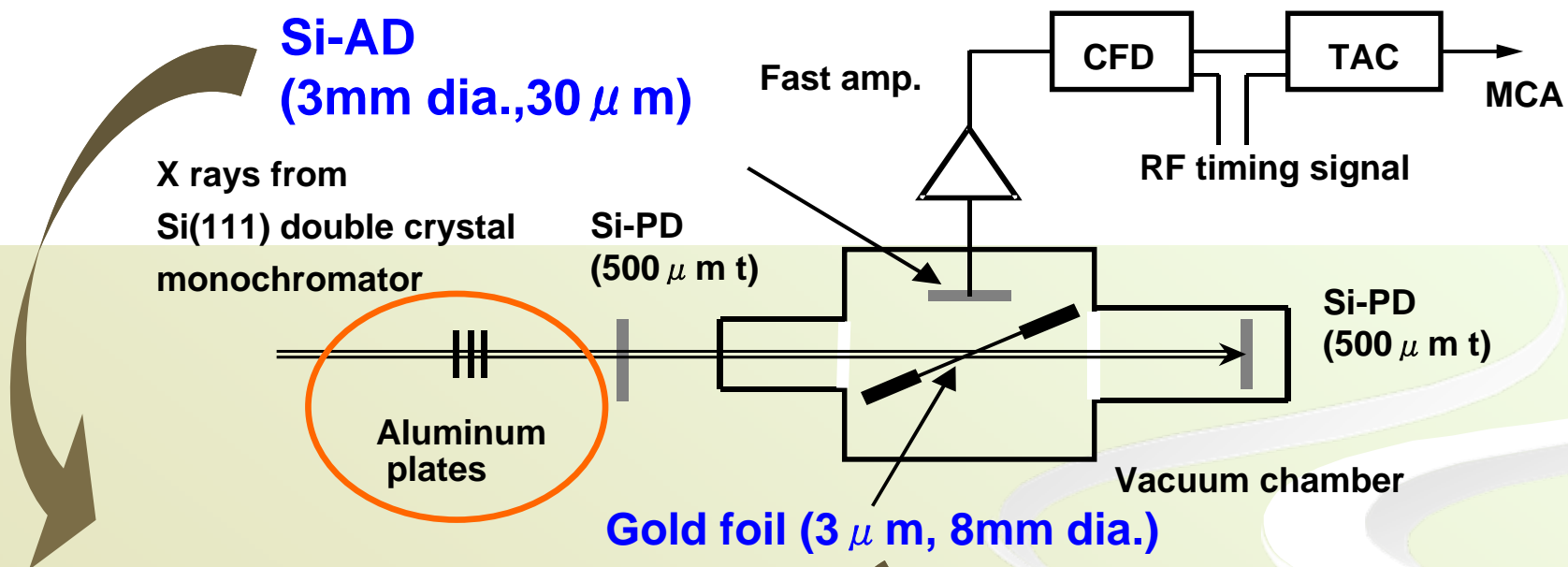


Fast response: suitable for time gating



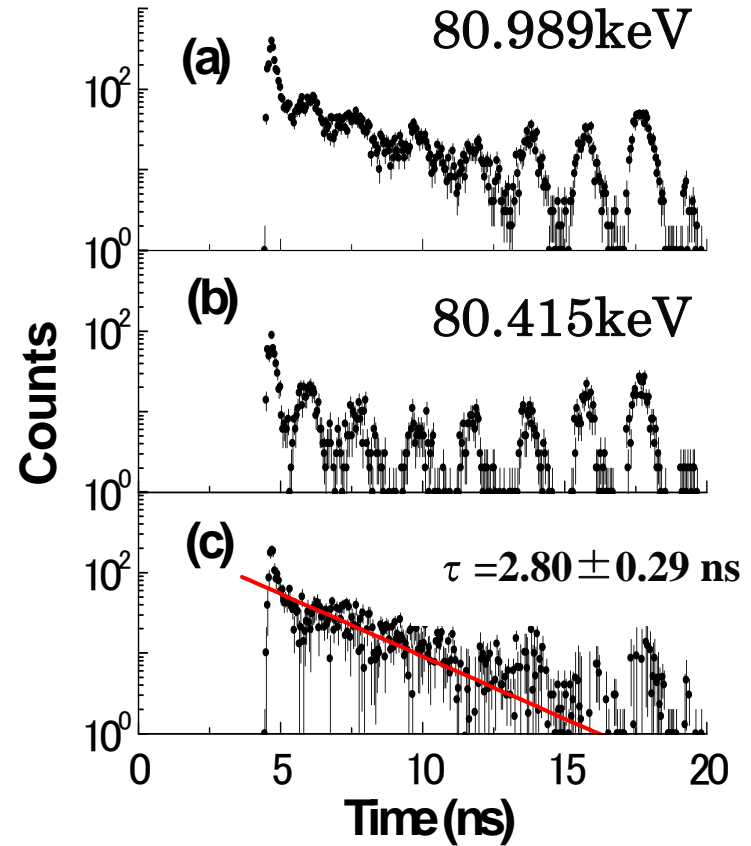
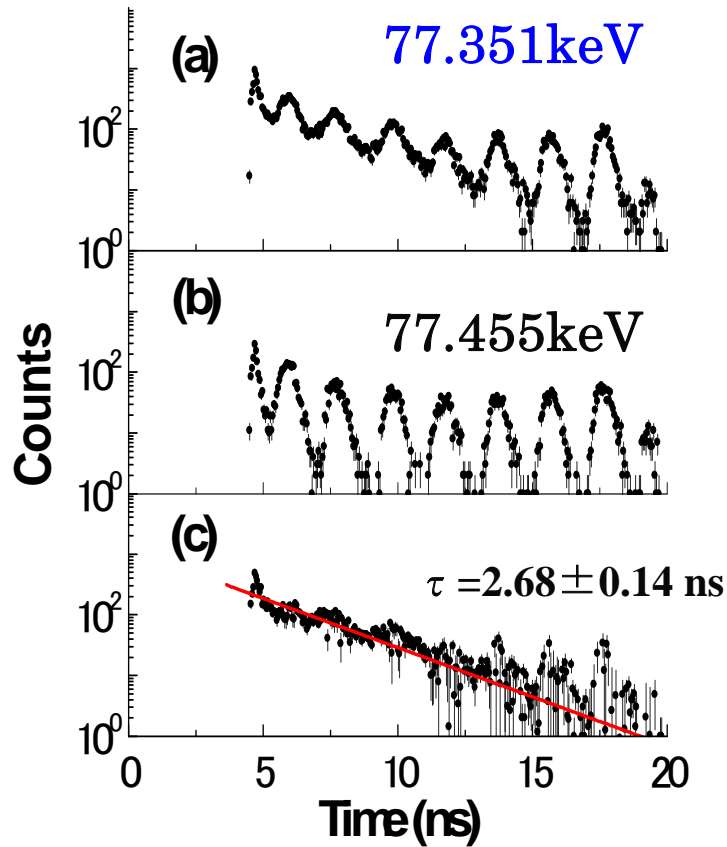
3. Experiments results on Au-197 and on Ir-193

NEET experiment on Au-197 at SPring-8 BL09XU in 1999 -



Results in Au-197 Exp.

Time spectra of Nuclear Resonance (left) & NEET (right)

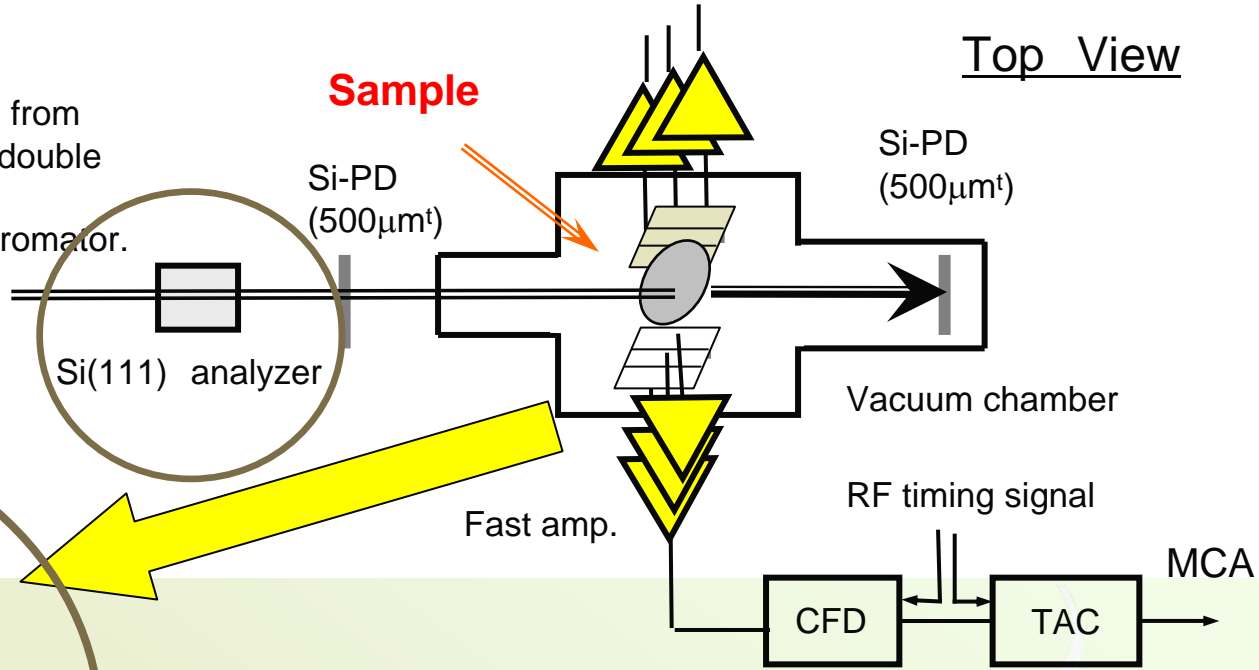


$$T_{1/2} = 1.91 \text{ ns} \rightarrow \tau = 2.76 \text{ ns}$$

$$P_{\text{NEET}} = (5.0 \pm 0.6) \times 10^{-8}$$

Improvements on the experimental set-up

x-rays from Si(333) double crystal monochromator.



Top View

1x6 mm², 3 ch, 30 μm^t

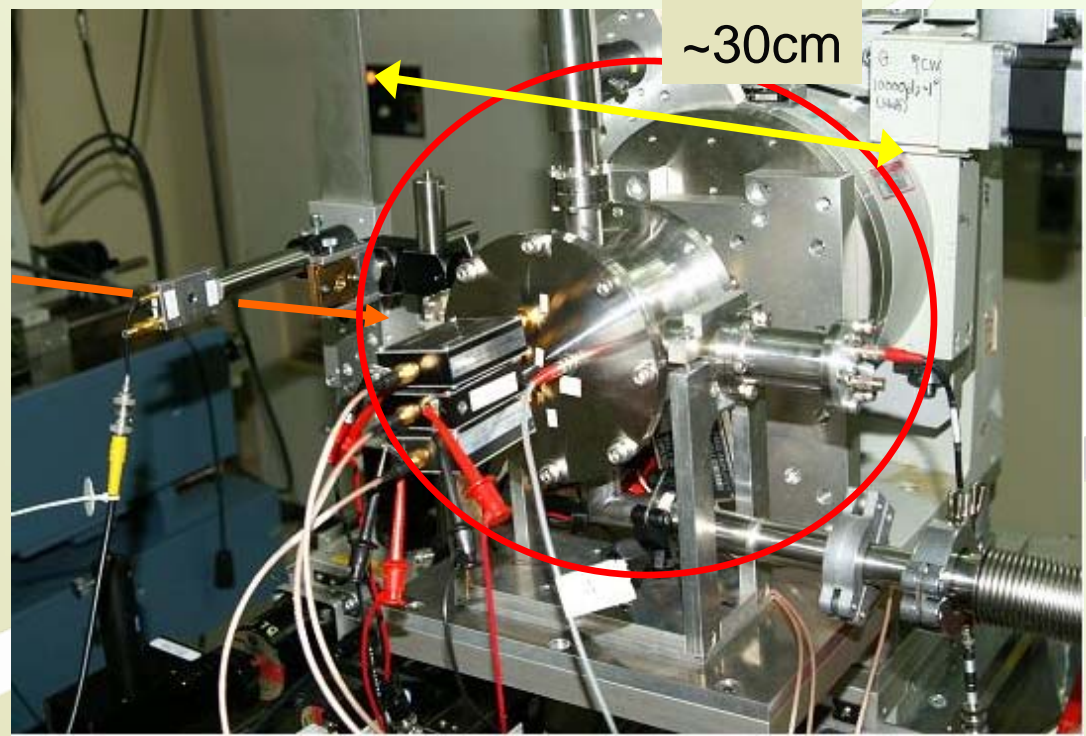


18 mm²

Φ 3mm, 30 μm^t

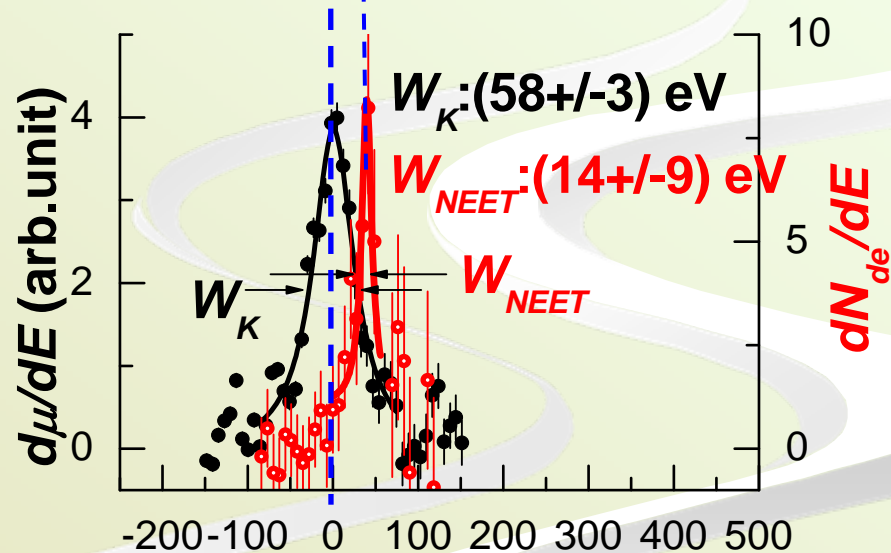
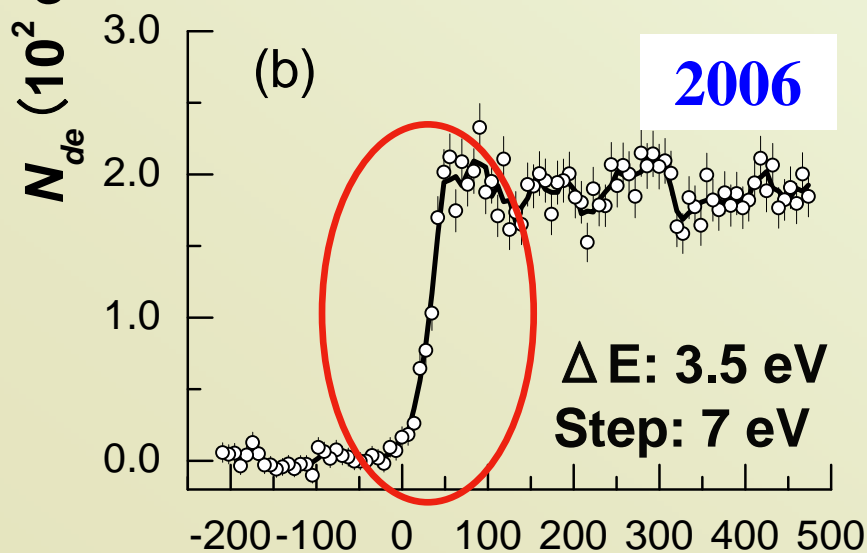
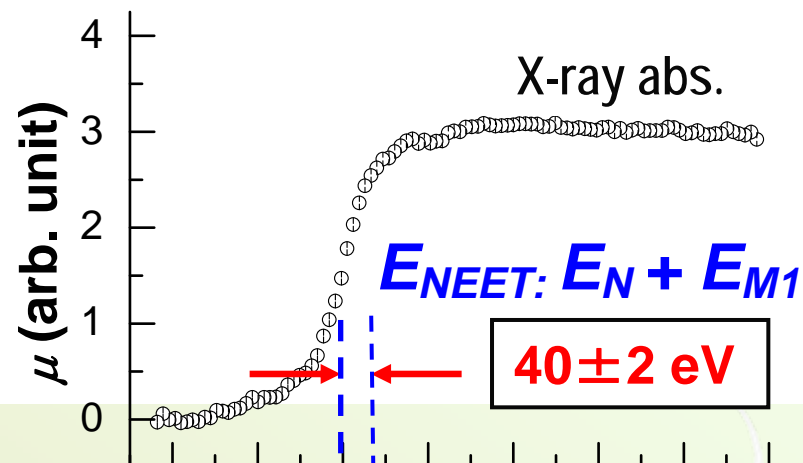
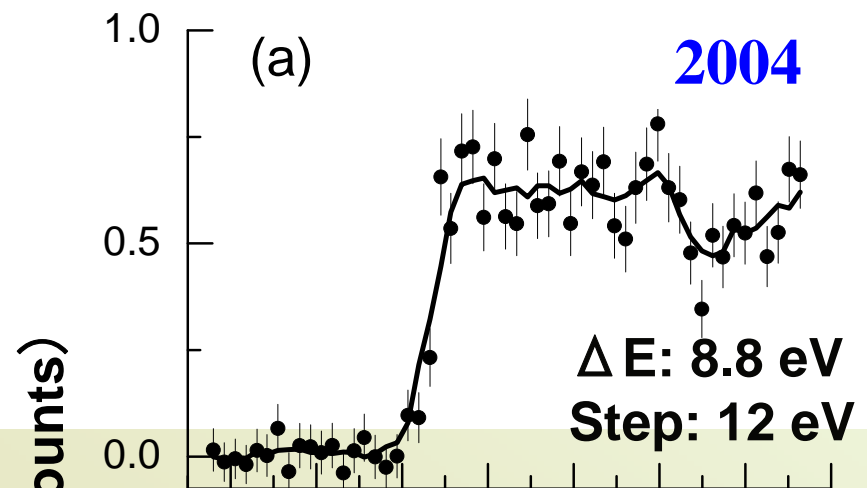


7 mm²



Nuclear events around the K -absorption edge

Phys.Rev.C74 (2006) 031301(R).



0.4 s $^{-1}$

ΔE_K (eV)

$\Delta E_K = 0$: K edge

ΔE_K (eV)

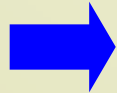
5. Next challenges

	E_{ion} (keV)	E_{M} (keV)	E_{N} (keV)	$T_{1/2}$ (ns)	ΔE (keV)	P_{NEET} (cal)
^{189}Os (Z=76) 16.1%	73.856 (K)	3.049(M ₁)	69.537	1.62	1.270	$^E 1.2 \times 10^{-10}$
^{187}Os (Z=76) 1.6%	12.971(L ₁)	3.049(M ₁)	*9.776 (9.746)	2.38	0.146	$^F 1 \times 10^{-8} ?$

^E E.A. Tkalya, Phys. Rev. A 75 (2007) 022509.

^F E. A. Tkalya, private communications.

A. ^{189}Os

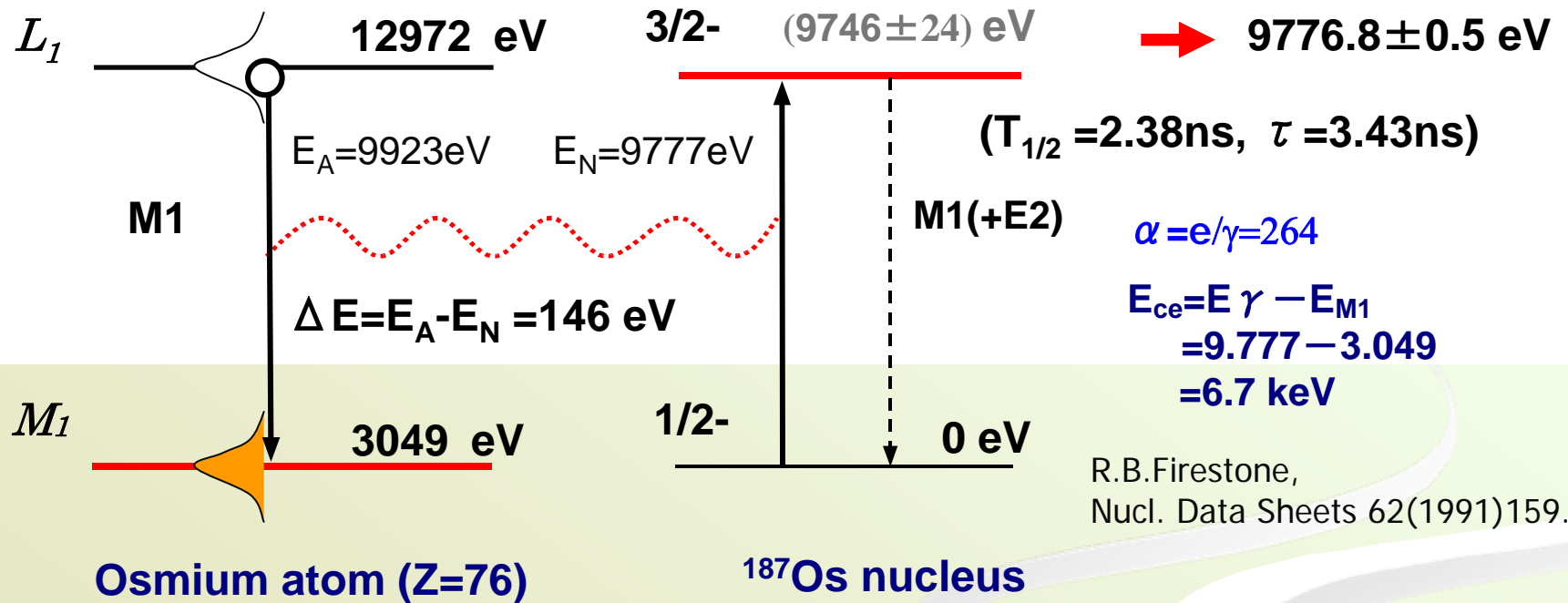


$\Delta E = 1.27 \text{ keV}$

P_{NEET} : very low ! < 1/10 of ^{193}Ir case

We need a larger efficiency in the detector system !

B. ^{187}Os

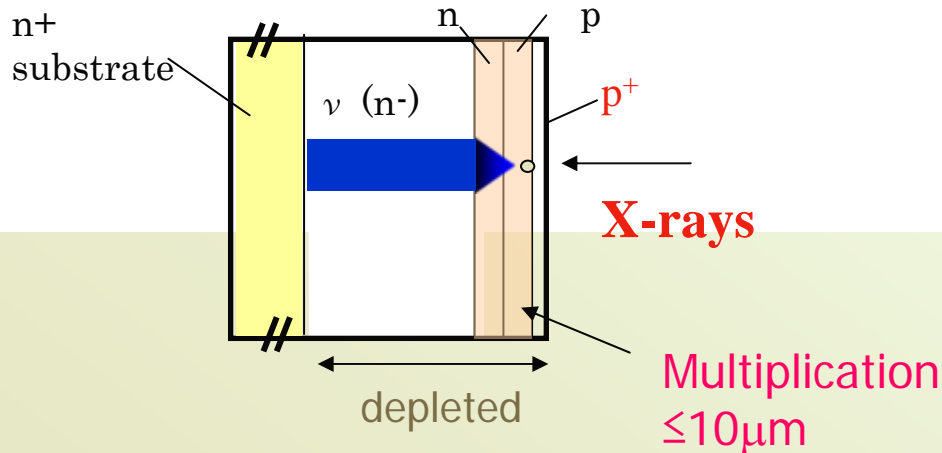


$1 \times 10^{-8} ?$: P_{NEET} may not be so small, relatively.

Problems are there in low energy of excitation X-rays (12.8 keV) and the M -internal conversion electrons (6.7 keV).

The detector has to be not sensitive for the elastic scattering X-rays, with keeping the efficiency for the int. conv. electrons.

Reverse Type of Si-AD – “radiation sensitive” only near the surface region



HM S8664 series

$C_p \sim 20\text{pF}$ (3mm dia.)

A thin ($< 5\text{-}\mu\text{m}$) junction to reduce the efficiency for elastic X-rays in detecting the low-energy electrons.

For Os-187 exp.,
7-keV Electron: $\lambda_e \sim 0.6 \mu\text{m}$

In a $5\text{-}\mu\text{m}$ layer,
12-keV X-rays: $\varepsilon \sim 2\%$
Cf. $30\text{-}\mu\text{m}$ dep. $\varepsilon \sim 12\%$

6. *Conclusions*

1. *NEET can be observed by the experiments using synchrotron X-rays and the Si-AD electron detector.*
2. *The NEET measurements on ^{197}Au and on ^{193}Ir are in progress of studying the fine structure in nuclear events from discussing existence of NEET.*
3. *Observing NEET on ^{189}Os and on ^{187}Os will be the next challenges with improving the efficiency and the device structure of the Si-AD detector.*